



# The source of Earth's water

(and other important things humans seem to like)

**March 1, 2025**

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LLNL-PRES-823251

A really big  
moon



Life



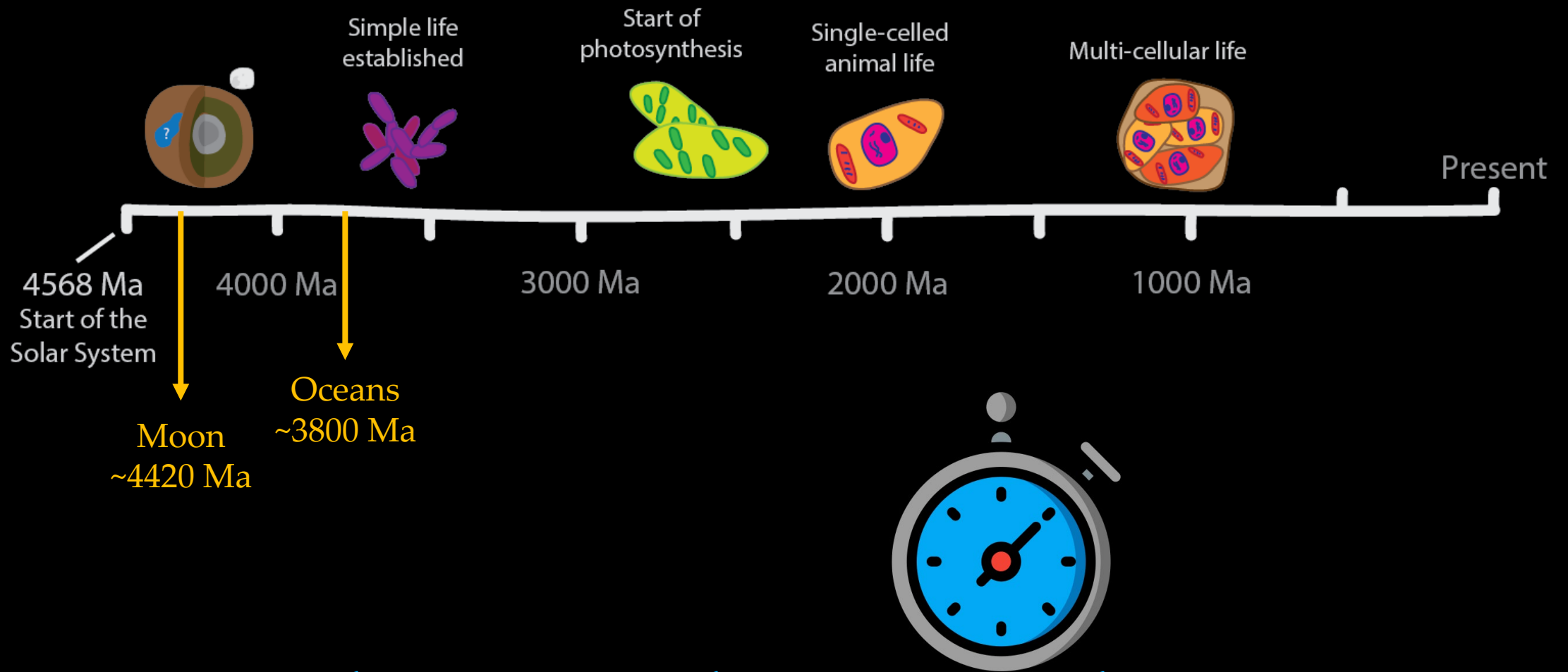
# What makes Earth “unique”?

(as far as we know)

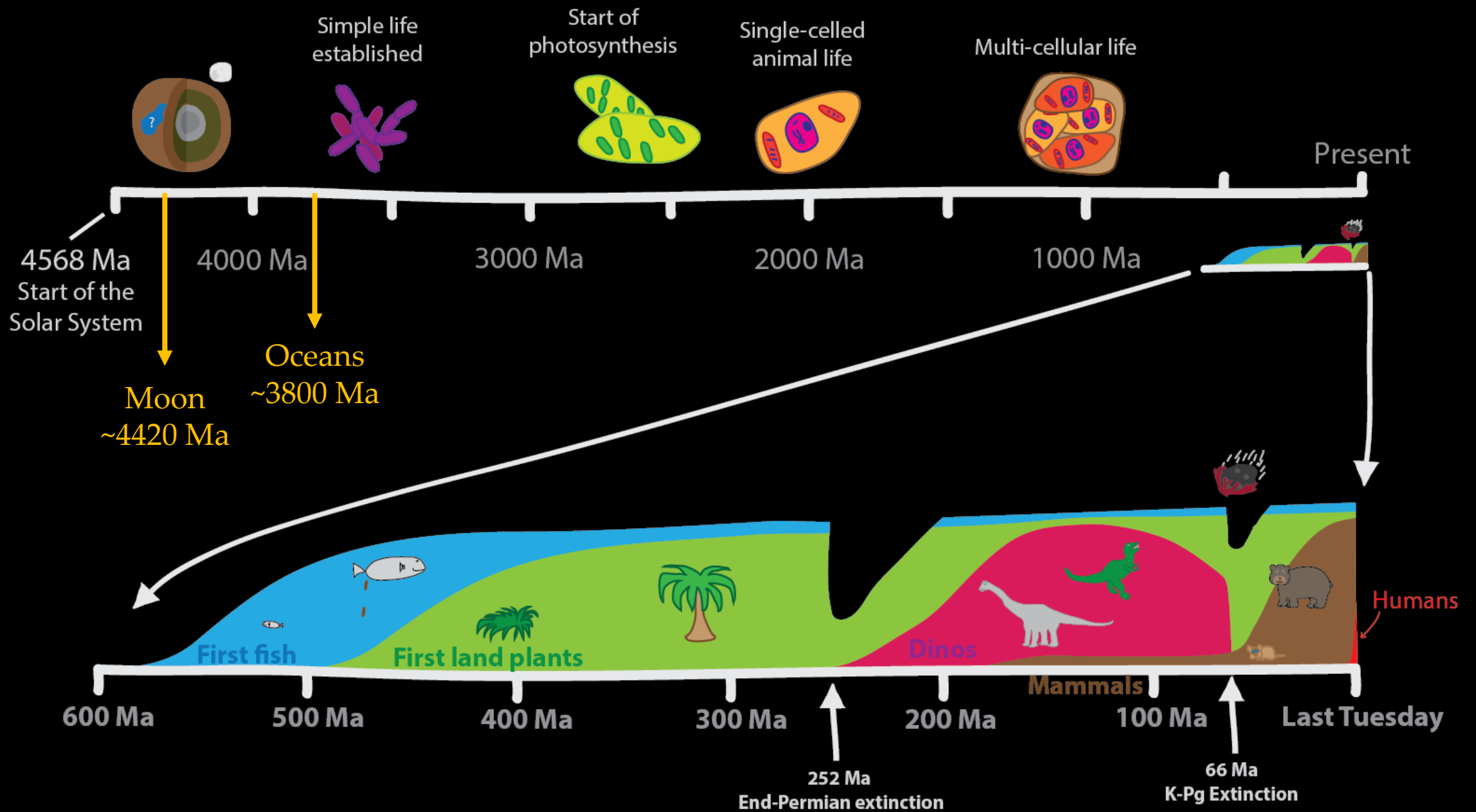
Lots of liquid water on the surface

iPhones

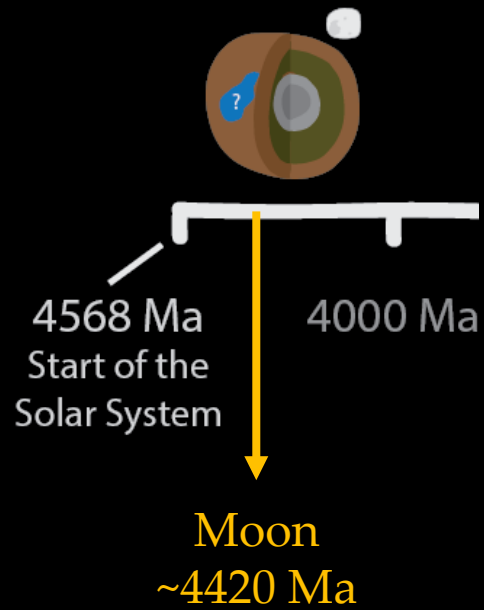




We know the timeline



# How did we get the Moon?






Collision

Debris disk

Happily  
ever after

# The secret ingredients are:

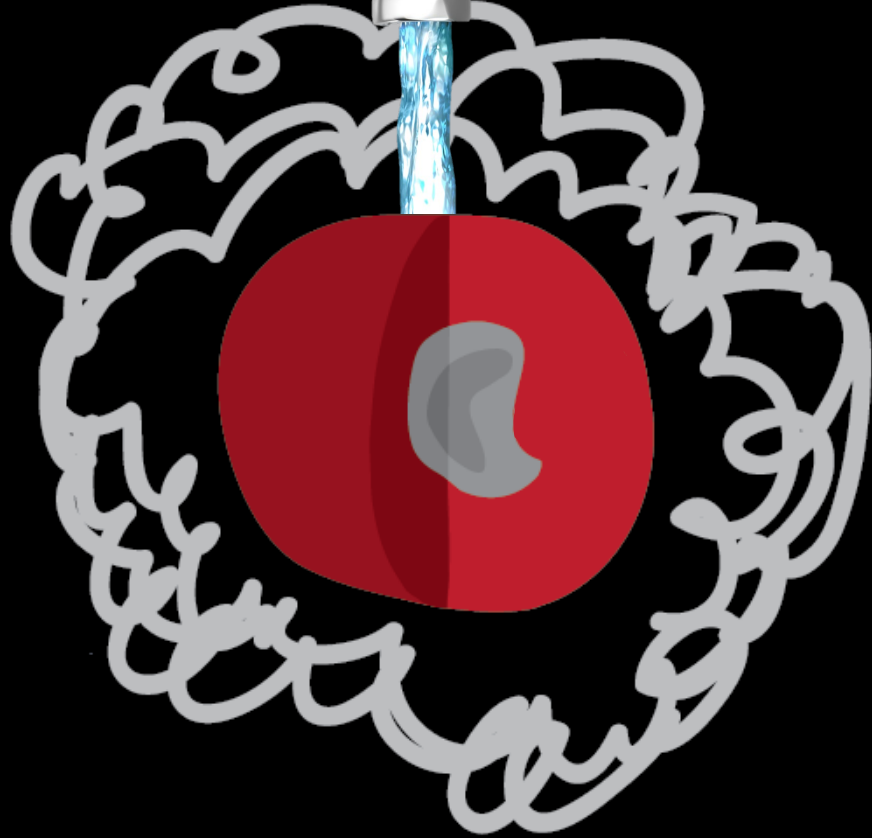
- *A really* big moon 
- Lots of liquid water on the surface
- Life
- iPhones



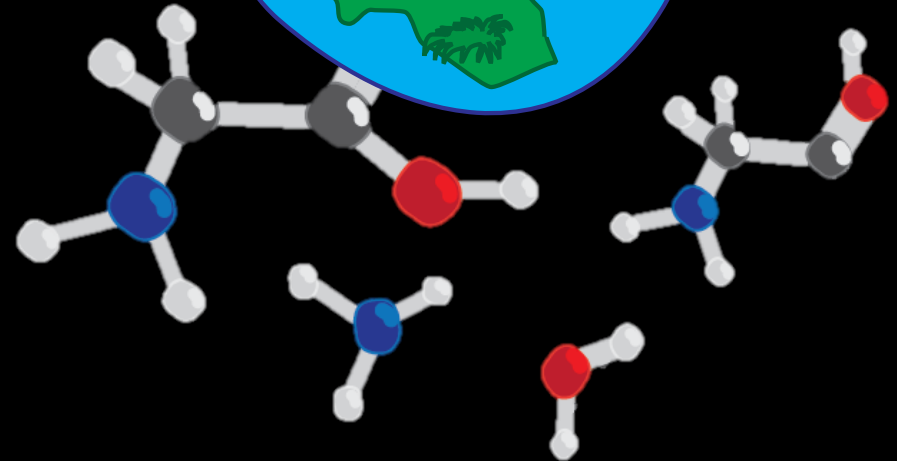
sterilized planet



**Water** (more on this later...)



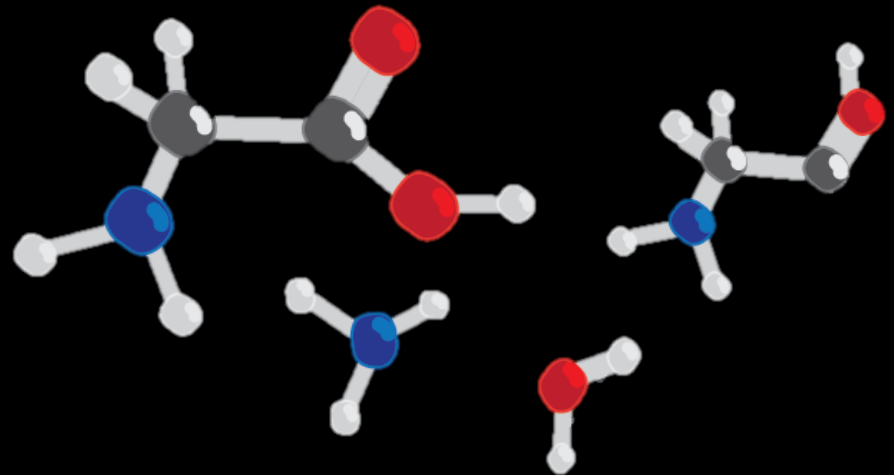
**Organic**  
molecules



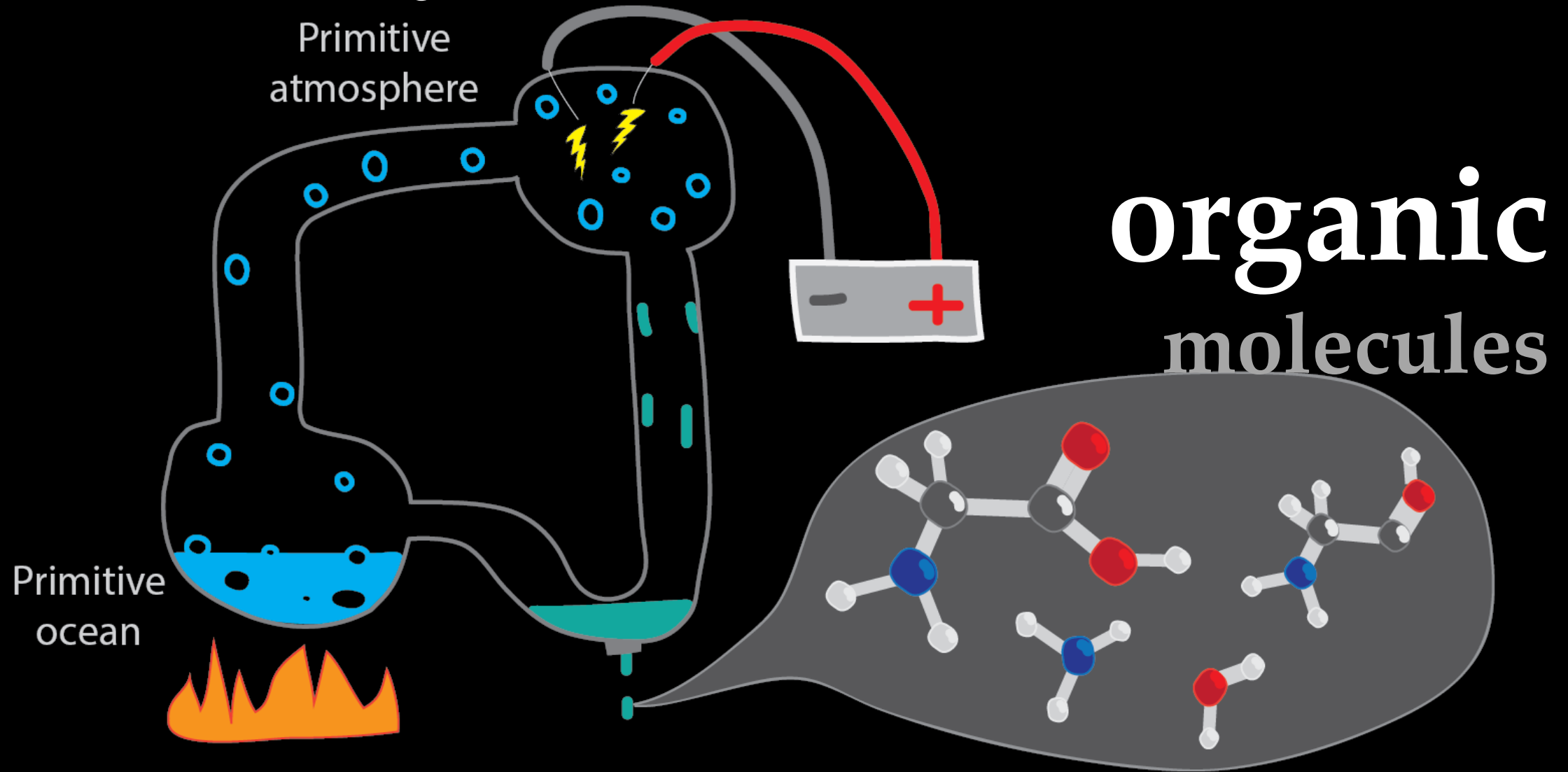
sterilized **planet**

# Possibility 1) produced on Earth

organic  
molecules

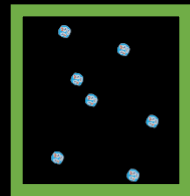


# Possibility 1) produced on Earth

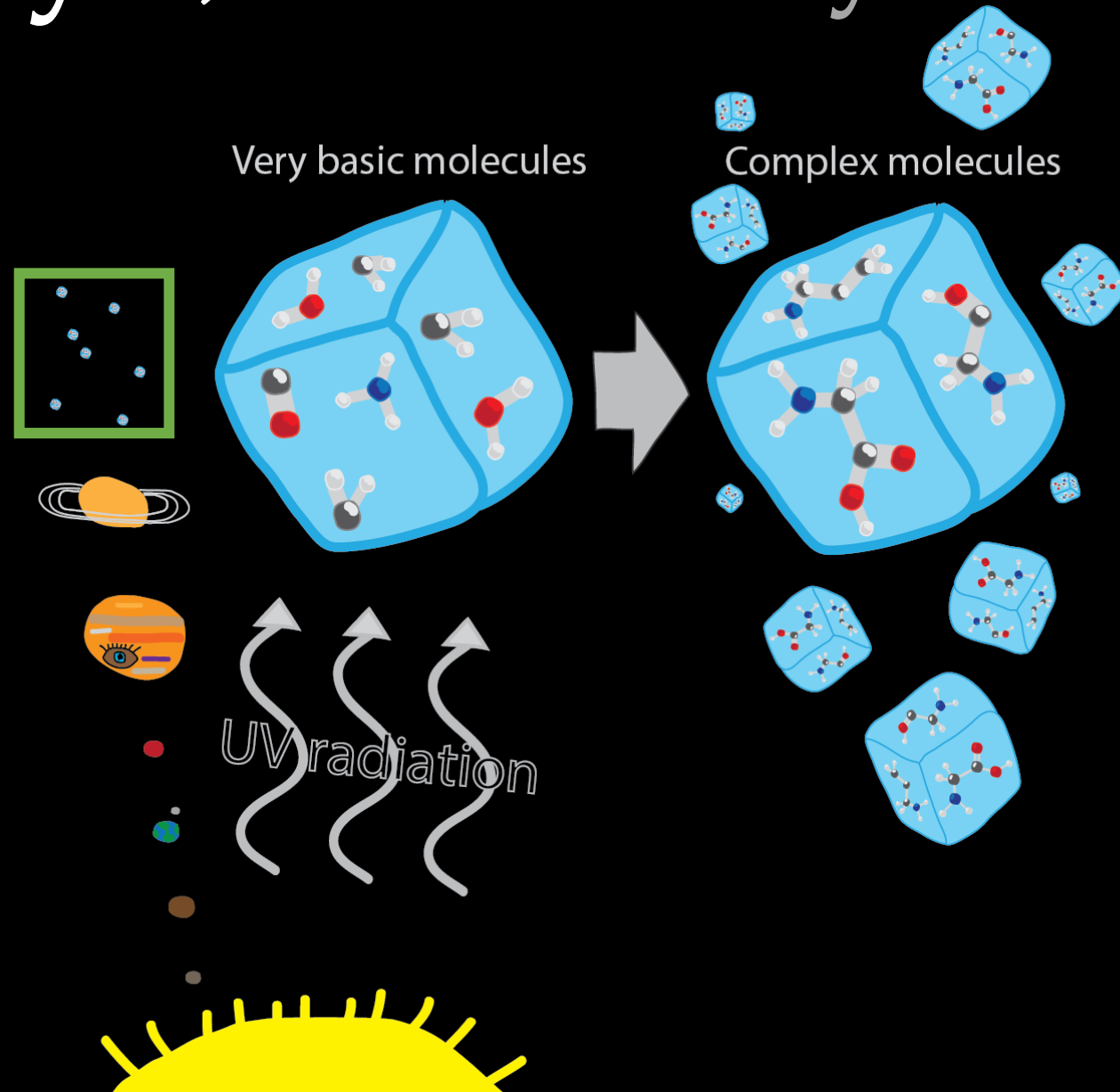


Possibility 2) delivered by meteorites

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## It is a numbers game...

Over 80 amino acids have been isolated  
Mannose, glucose, acetone, methanol, benzene...

3 of the 4 RNA nucleotide bases

2 of the 4 DNA nucleotide bases

All in 1 meteorite

UV radiation



Murchison meteorite



# Possibility 2) delivered by meteorites

## It is a numbers game...

4+ billion years ago, Earth was receiving  
~275 metric tons of carbon  
compounds...

## PER DAY



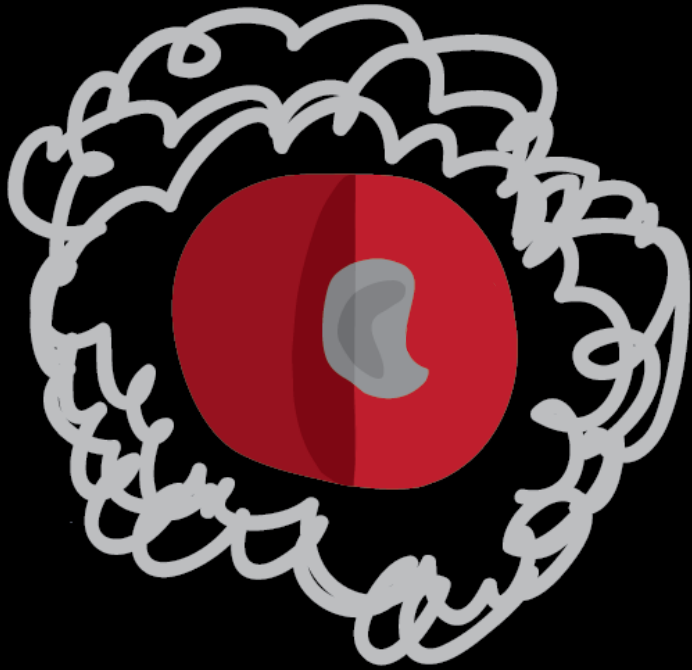
Murchison meteorite

# The secret ingredients are:

- *A really* big moon
- Lots of liquid water on the surface
- Life
- iPhones



# accretion of raw materials



Molten Earth

|  |   |  |   |
|--|---|--|---|
| 78<br>3827<br>1772<br><b>Pt</b><br>[Xe]4f <sup>14</sup> 5d <sup>9</sup> 6s<br>21.4 2.4 | 77<br>4428<br>2443<br><b>Ir</b><br>[Xe]4f <sup>14</sup> 5d <sup>7</sup> 6s <sup>2</sup><br>22.4 2,3,4,6 | 79<br>2857<br>1064.58<br><b>Au</b><br>[Xe]4f <sup>14</sup> 5d <sup>10</sup> 6s<br>18.9 1,3 | 45<br>3697<br>1963<br><b>Rh</b><br>[Kr]4d <sup>8</sup> 5s<br>12.4 2,3,4 |
|--|---|--|---|

|  |   |  |   |
|--|---|--|---|
| 78<br>3827<br>1772<br><b>Pt</b><br>[Xe]4f <sup>14</sup> 5d <sup>9</sup> 6s<br>21.4 2.4 | 77<br>4428<br>2443<br><b>Ir</b><br>[Xe]4f <sup>14</sup> 5d <sup>7</sup> 6s <sup>2</sup><br>22.4 2,3,4,6 | 79<br>2857<br>1064.58<br><b>Au</b><br>[Xe]4f <sup>14</sup> 5d <sup>10</sup> 6s<br>18.9 1,3 | 45<br>3697<br>1963<br><b>Rh</b><br>[Kr]4d <sup>8</sup> 5s<br>12.4 2,3,4 |
|--|---|--|---|



# The secret ingredients are:

- *A really* big moon
- **Lots of liquid water on the surface**
- Life
- iPhones



# Does Earth really have *that* much water?

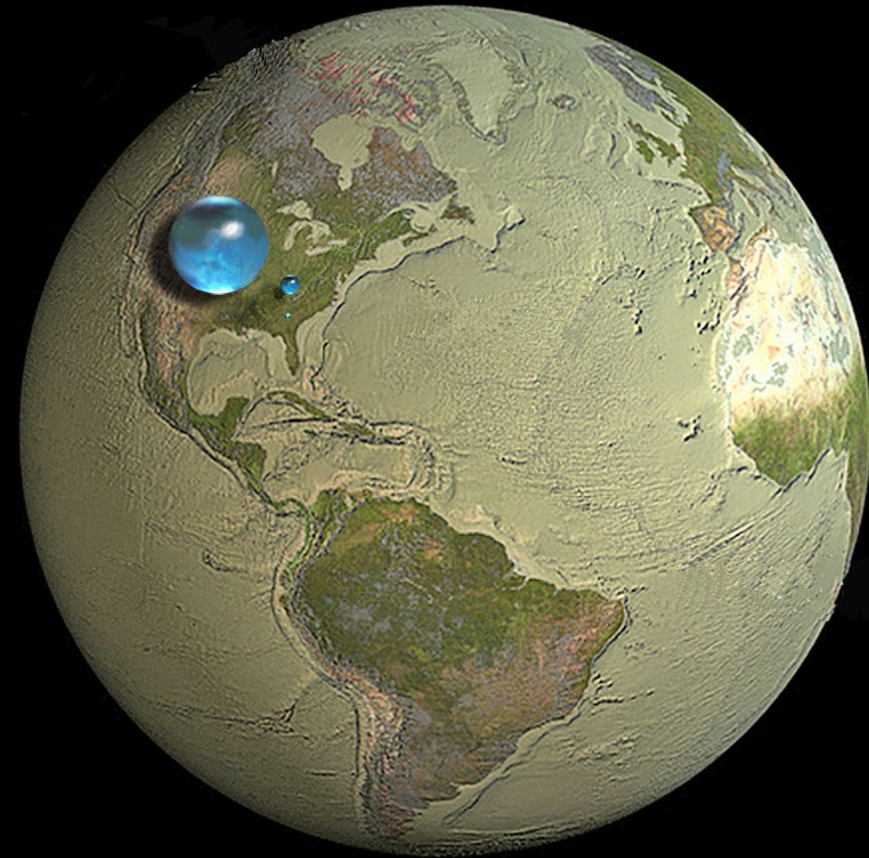
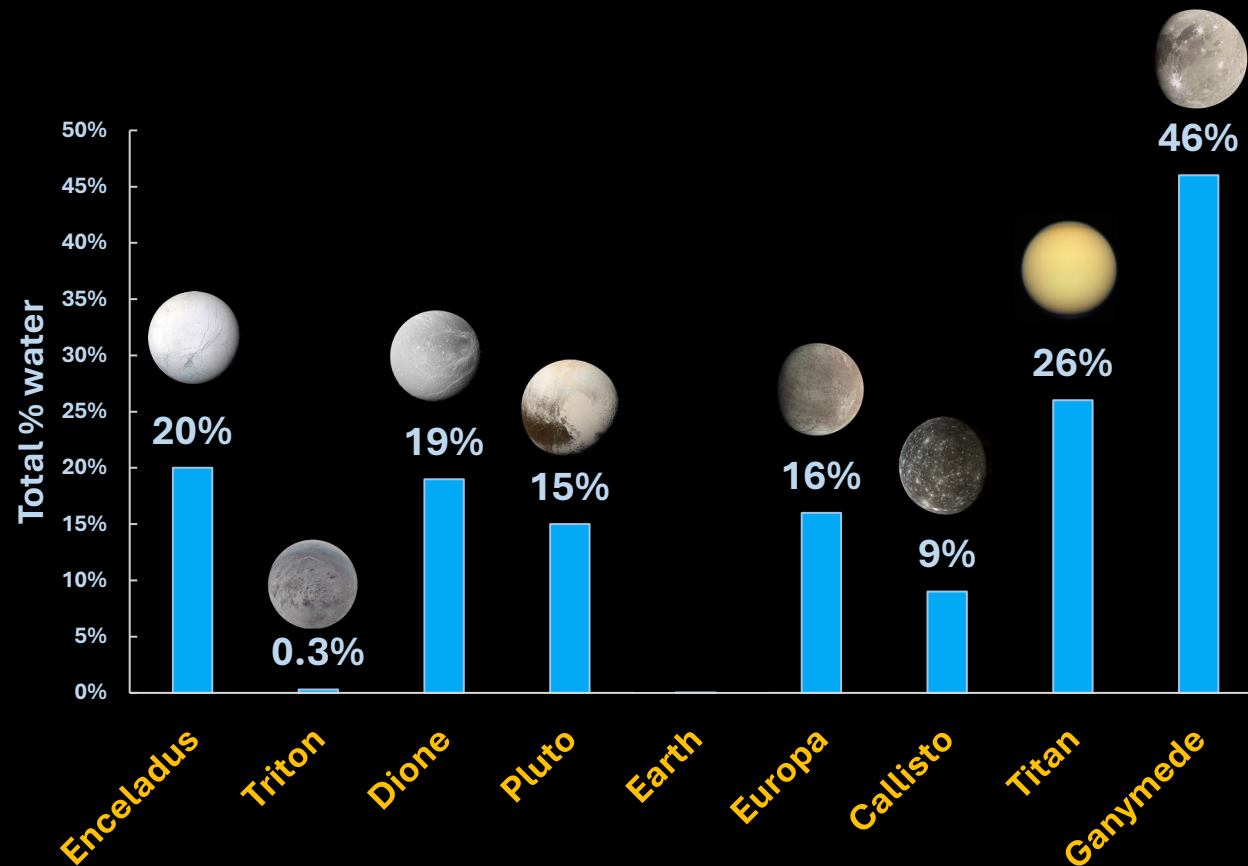


Image credit: Jack Cook, WHOI

**Total liquid water in various Solar System objects**  
(total by volume and water ice is not included)

Compared to our neighbors

# Original H<sub>2</sub>O

Lack of a magnetic field has led to major changes



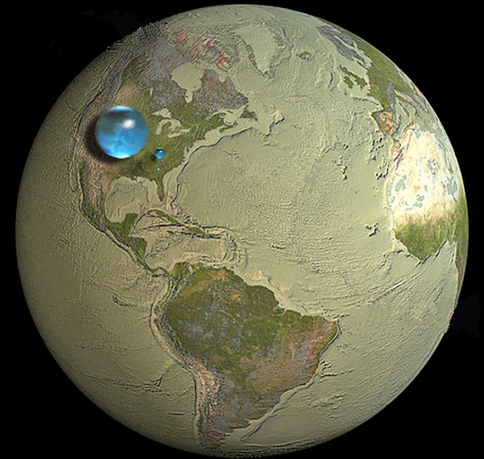
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But it does have water, and...

- It is on the surface
- It is liquid



What is the  
*source* of  
Earth's water?

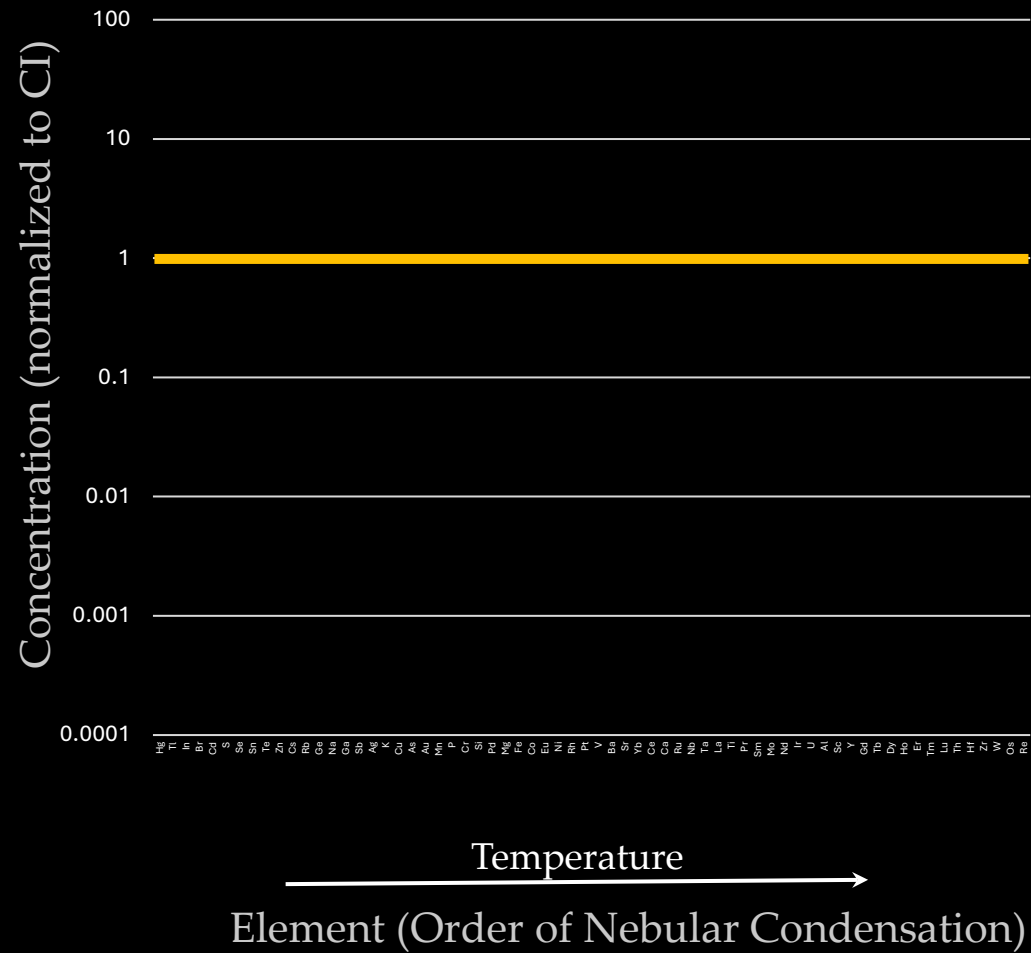


# There aren't that many options:

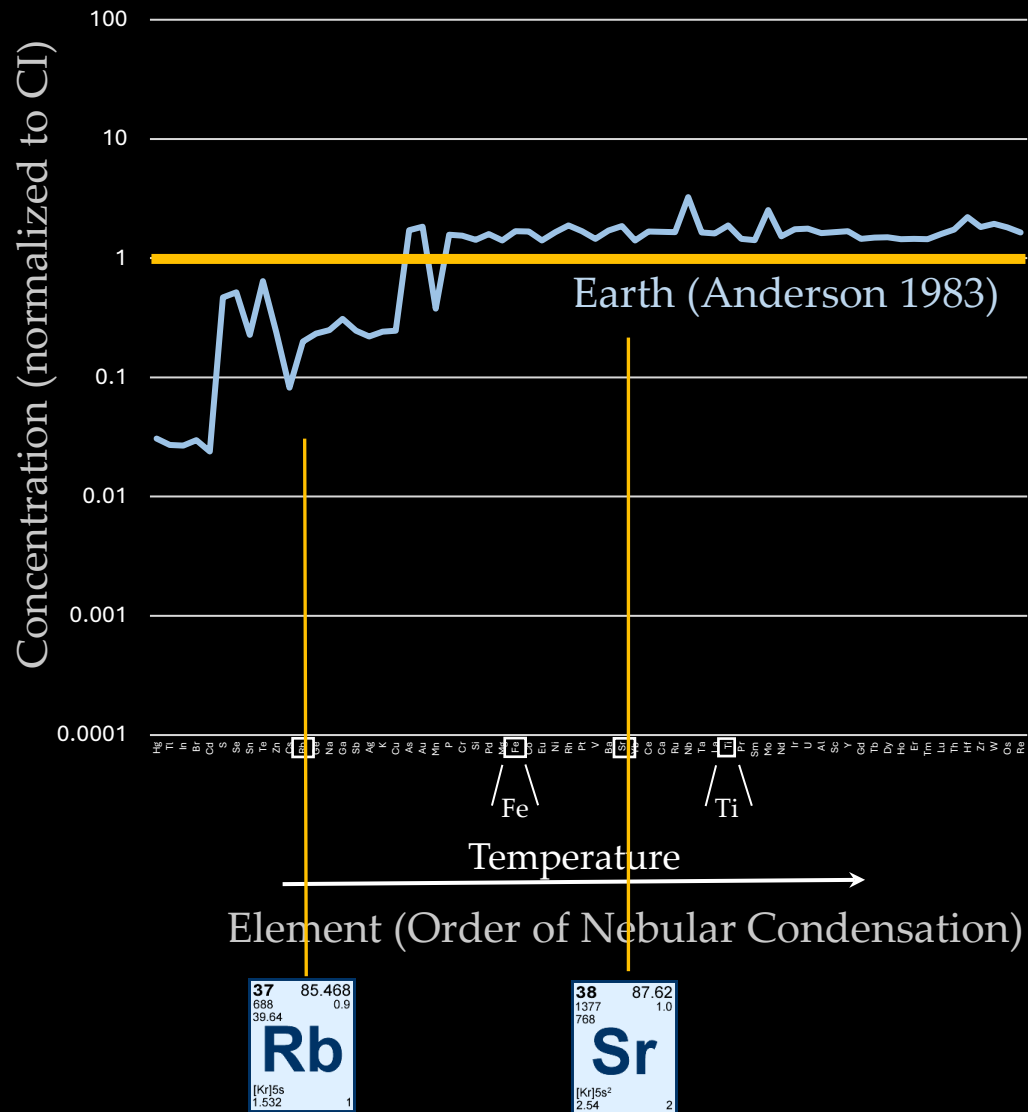
1. It was here from the start
2. It was delivered here
  - By meteorites
  - By comets



# Approach – Using the Moon

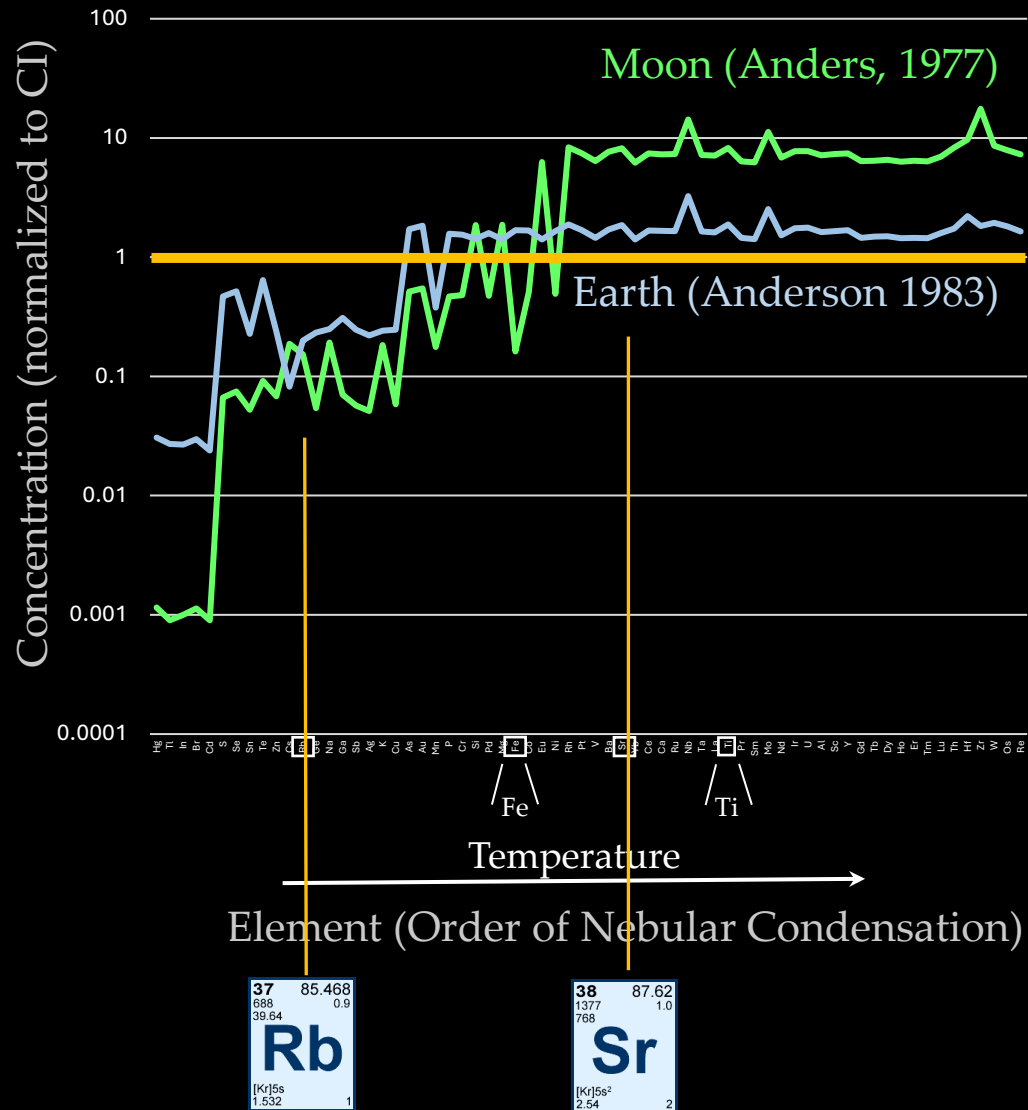


# Approach – Using the Moon



The Earth is depleted in  
volatile elements relative to  
the bulk Solar System

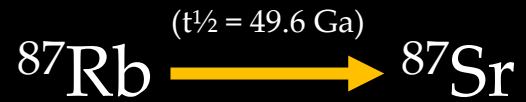
# Approach – Using the Moon



The Earth is depleted in volatile elements relative to the bulk Solar System

The Moon is more depleted than the Earth

# Approach

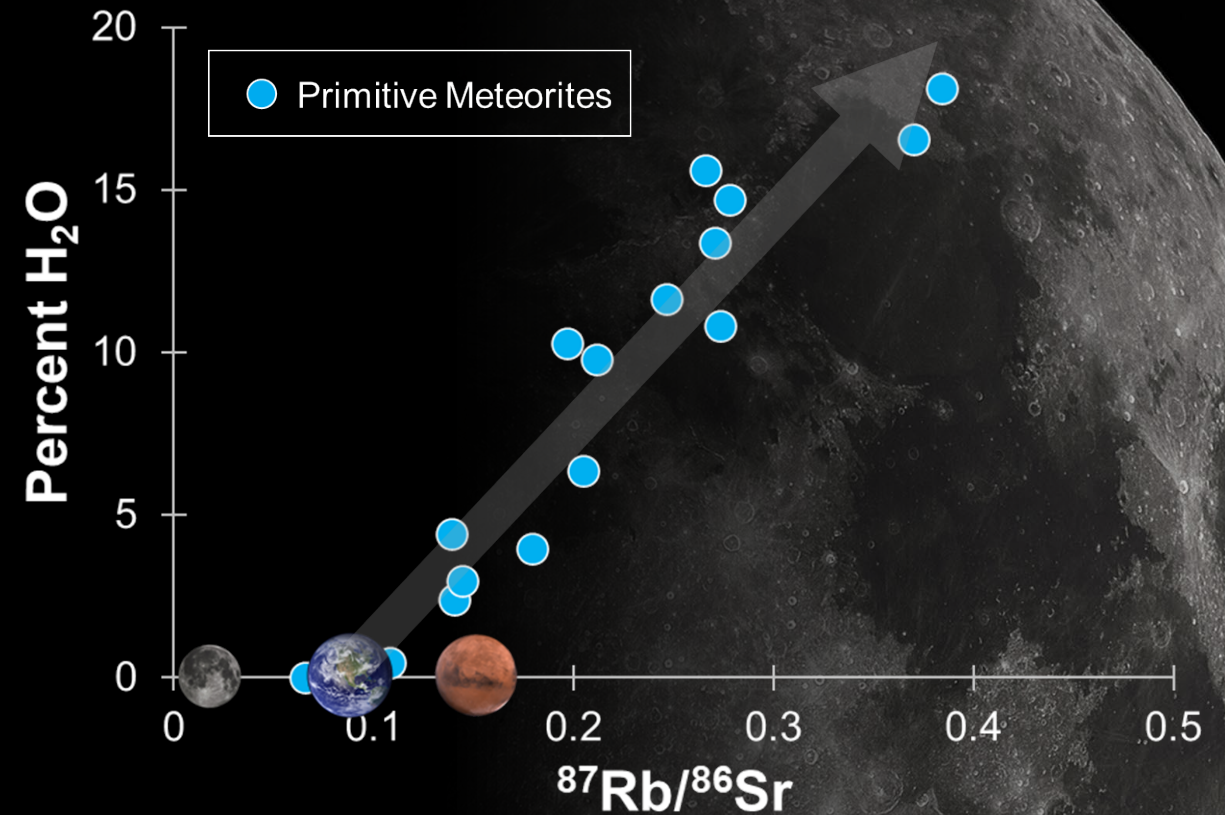


- Rb is a moderately volatile element
- Sr is refractory element

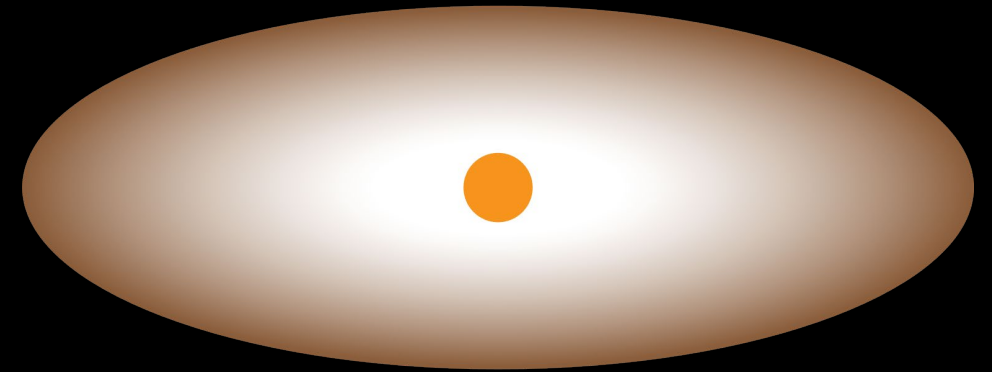
The  $^{87}\text{Sr}$  in material records time-integrated  $^{87}\text{Rb}$  abundance

The time-integrated  $^{87}\text{Rb}$  abundance can be calculated from the known  $^{87}\text{Sr}$  abundance of a material today

Rb is a proxy for  $\text{H}_2\text{O}$



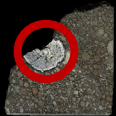
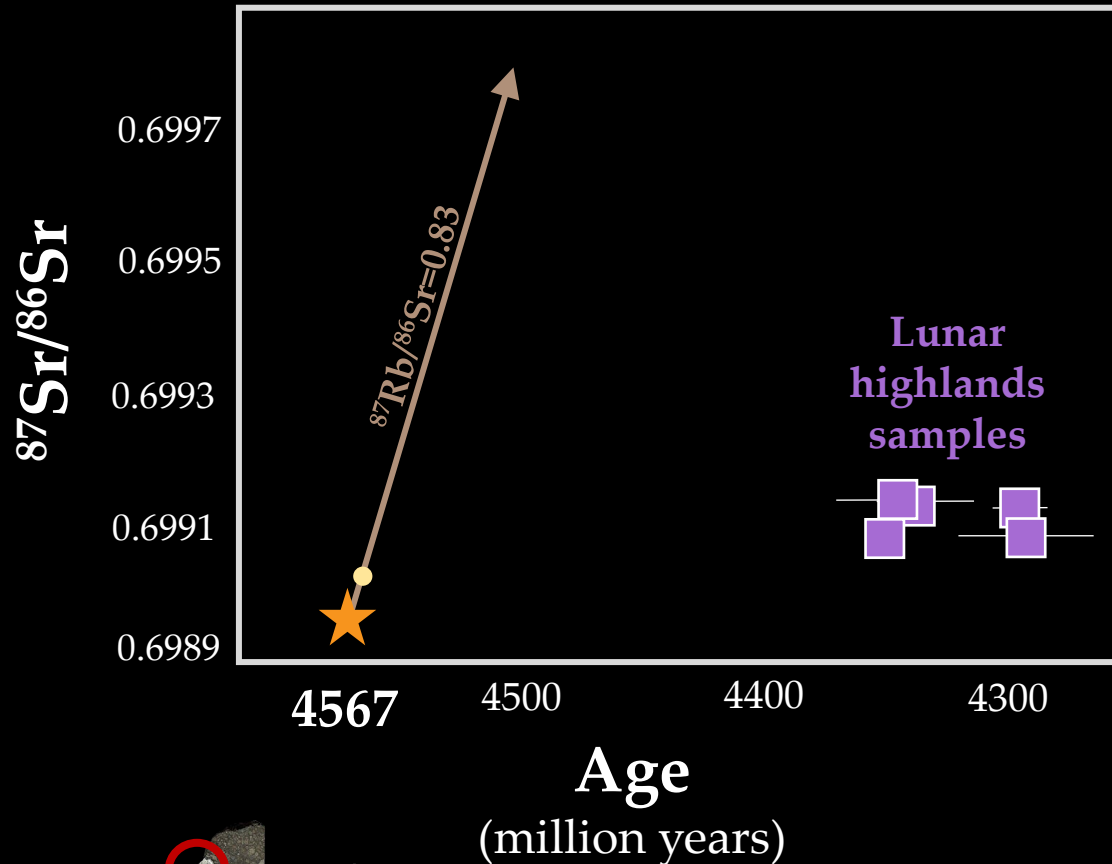
## Stage 1: Protoplanetary Disk

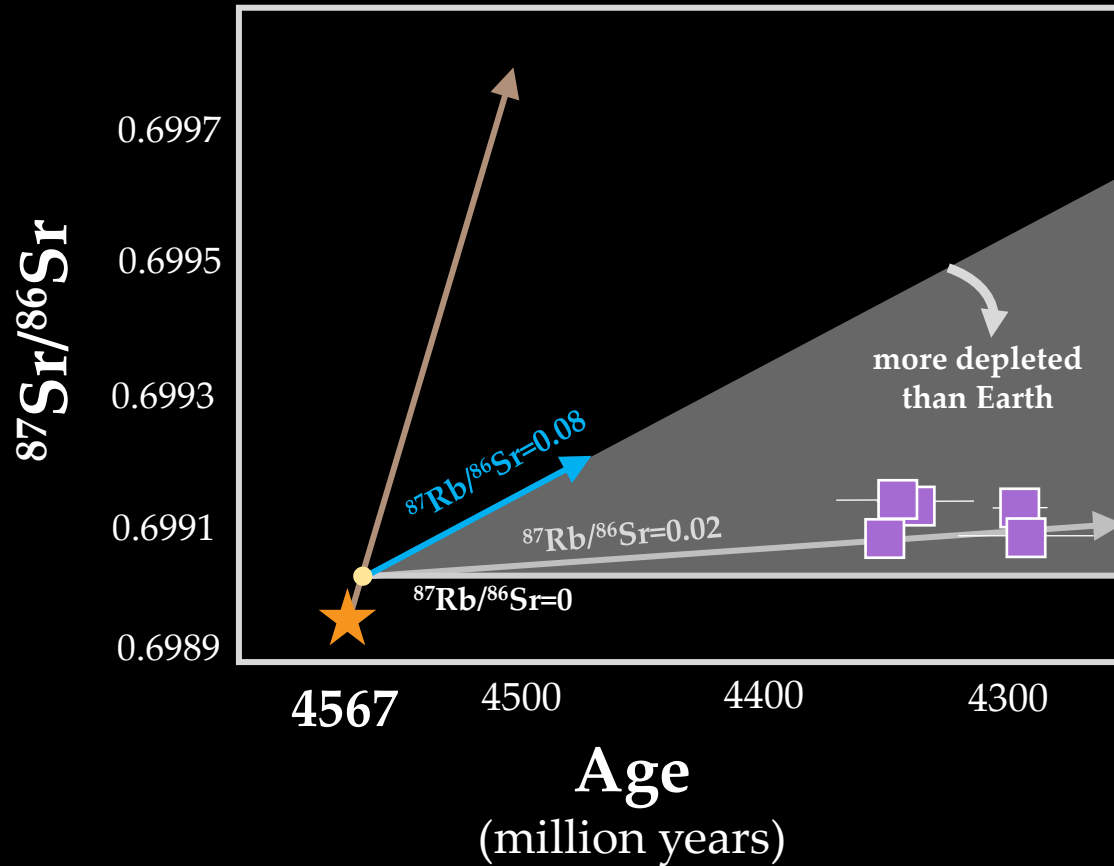


High Rb/Sr ( $\sim 0.83$ )

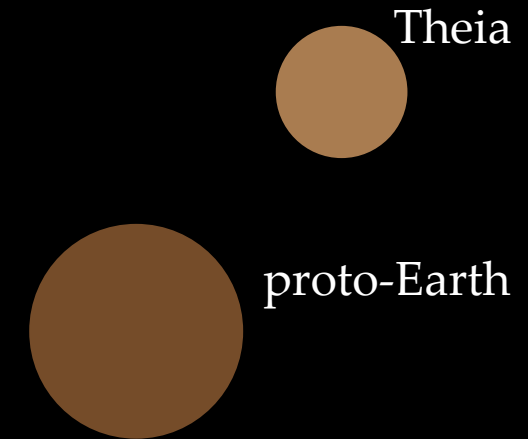
Grows in a lot of  $^{87}\text{Sr}$

Only lasts 1-3 million years





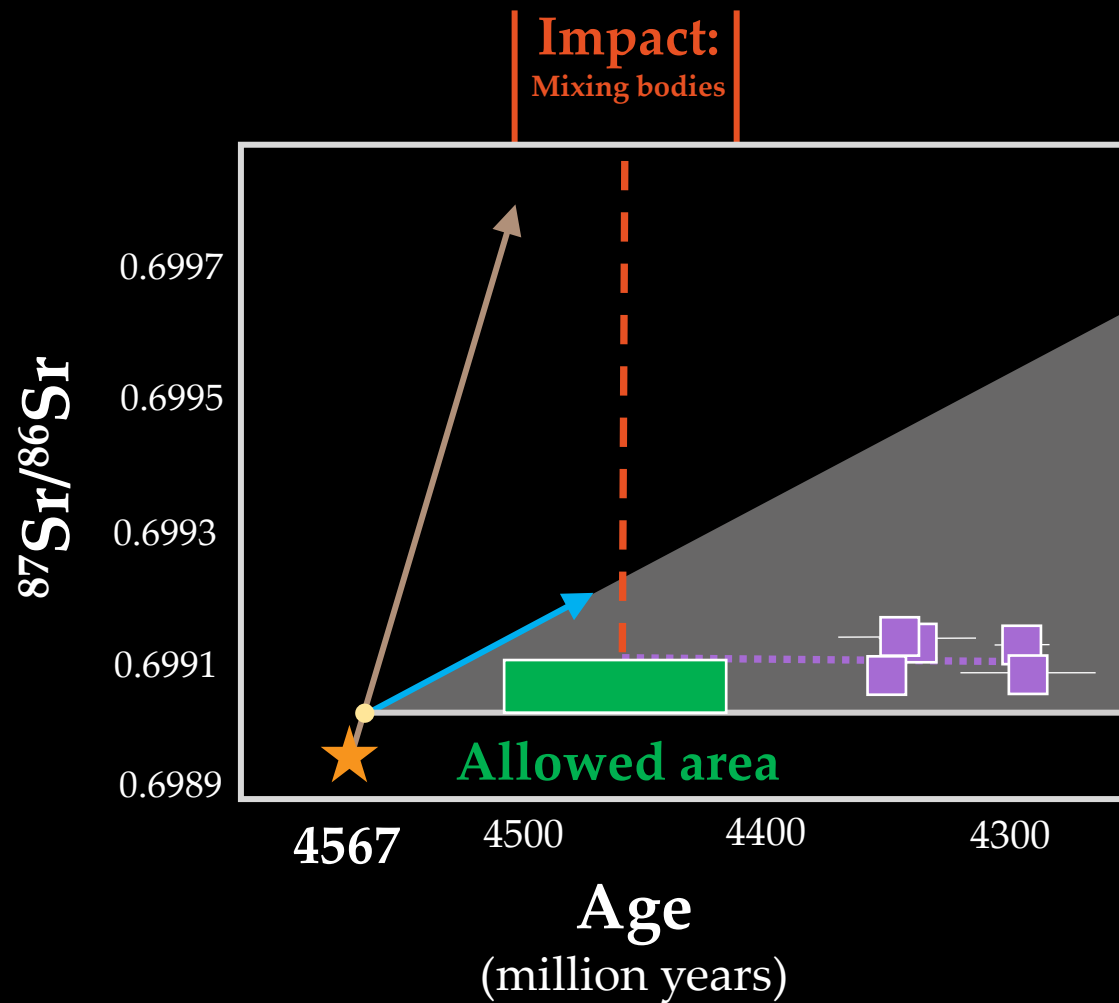
## Stage 2: Precursor Bodies



**We don't know the  $^{87}\text{Rb}/^{86}\text{Sr}$  (volatiles)  
...but we can set constraints!**

**We know the  $^{87}\text{Rb}/^{86}\text{Sr}$  of bulk Earth**

**We know the  $^{87}\text{Rb}/^{86}\text{Sr}$  of bulk Moon**

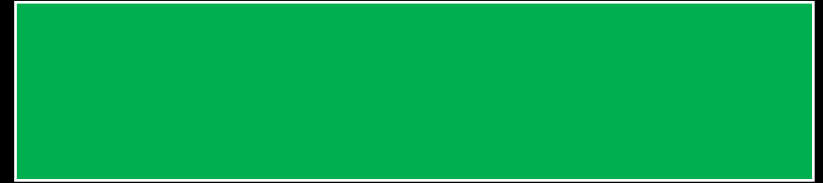
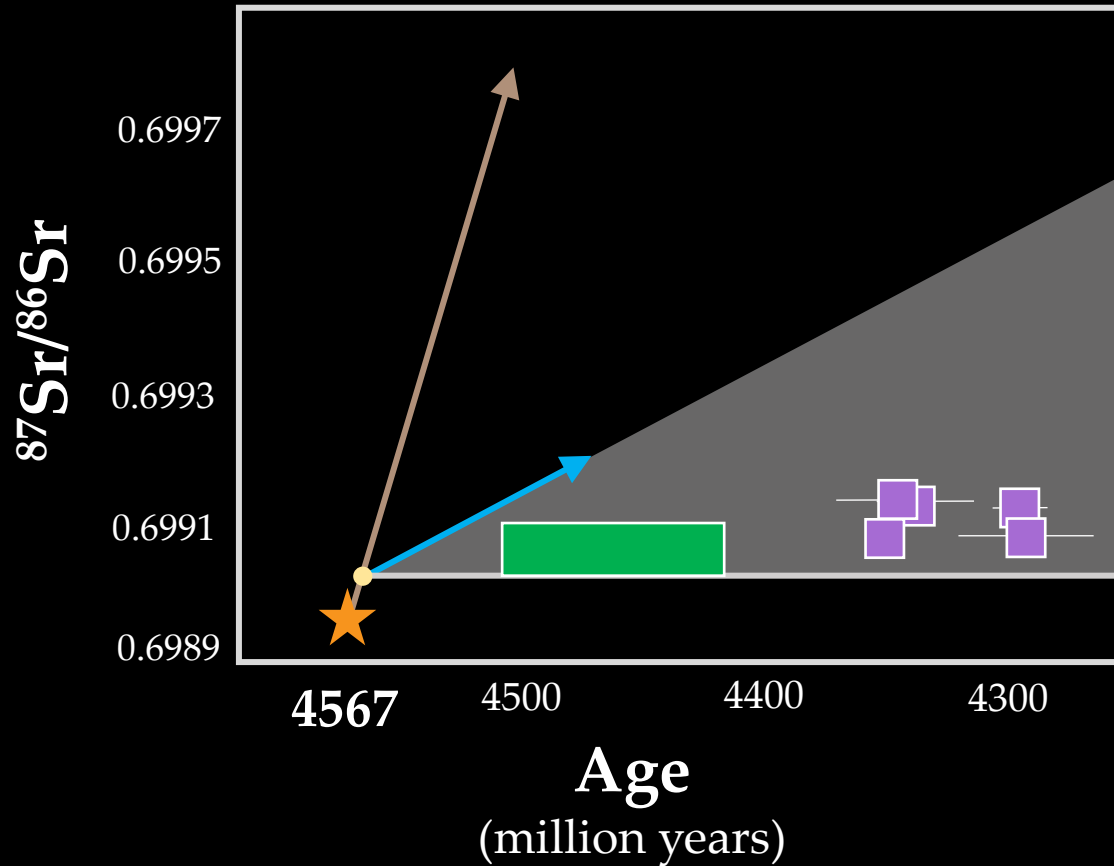


### Stage 3: **GIANT IMPACT!** Undifferentiated Moon

Earth & debris disk

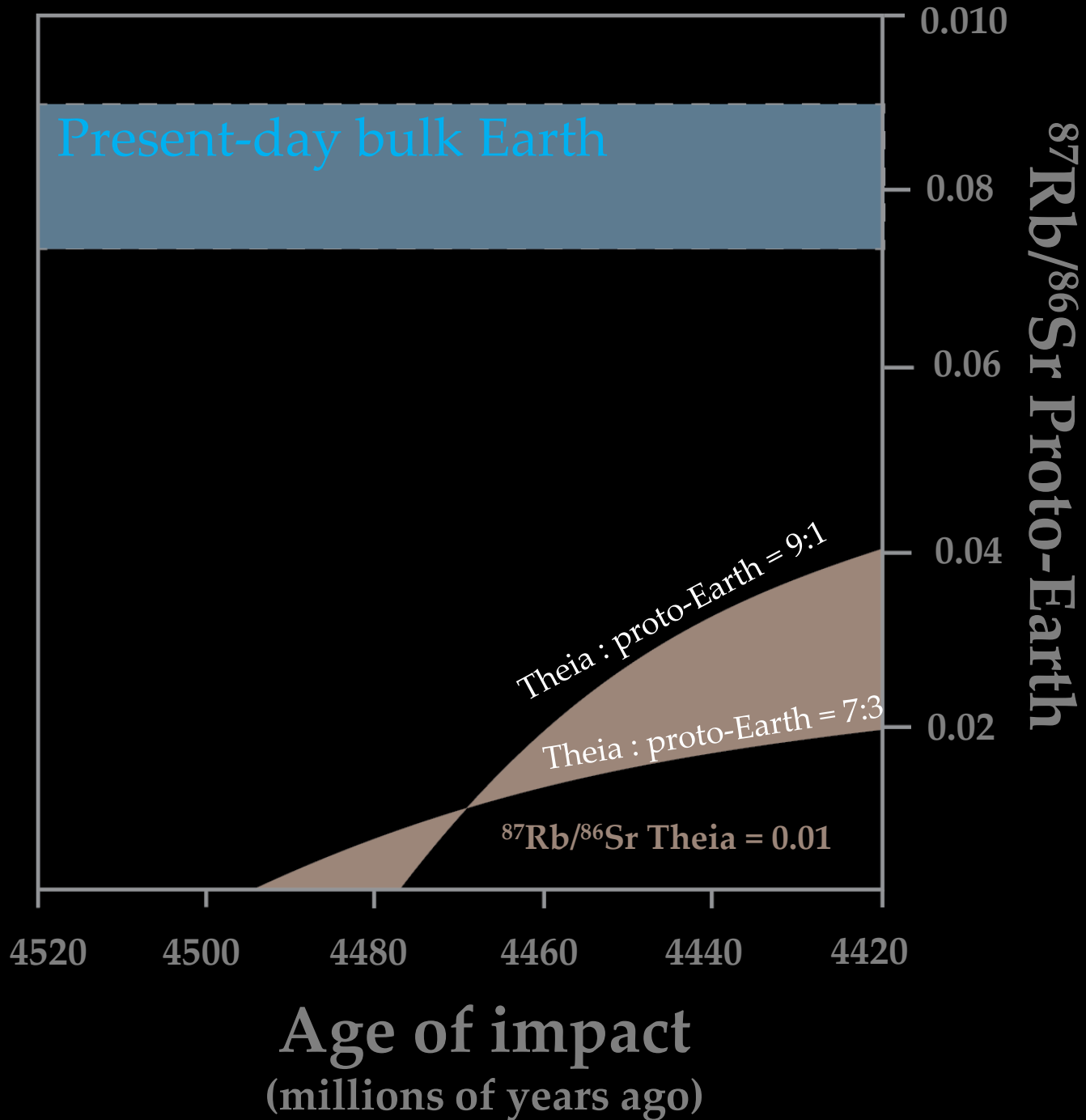
Earth-Moon system



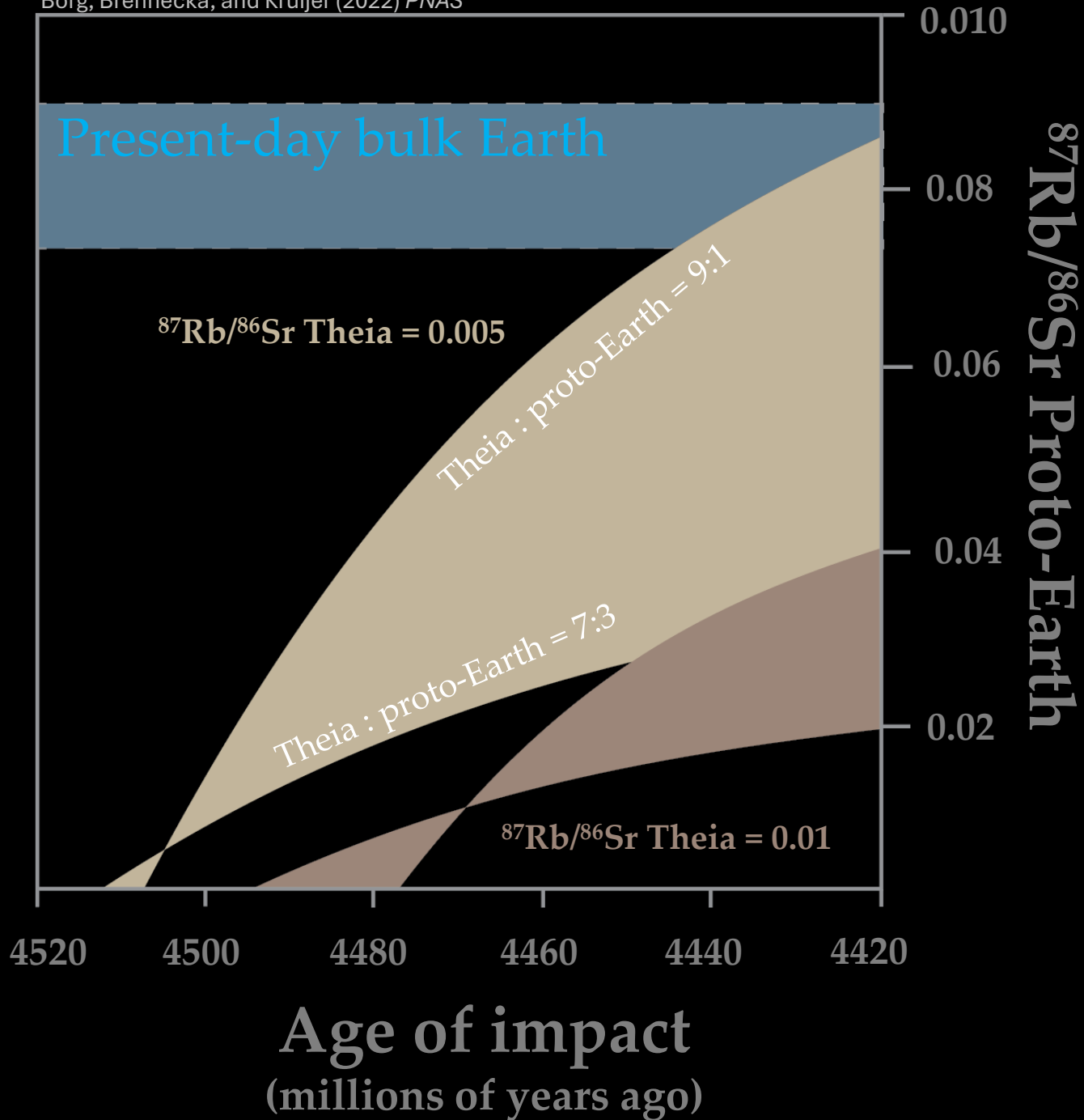


This box sets hard  
constraints on:

1. When the Moon-forming impact happened
2. The volatile content of both Earth and Theia



With an  $^{87}\text{Rb}/^{86}\text{Sr} = 0.01$  for  
Theia, you cannot  
reproduce present-day  
Earth

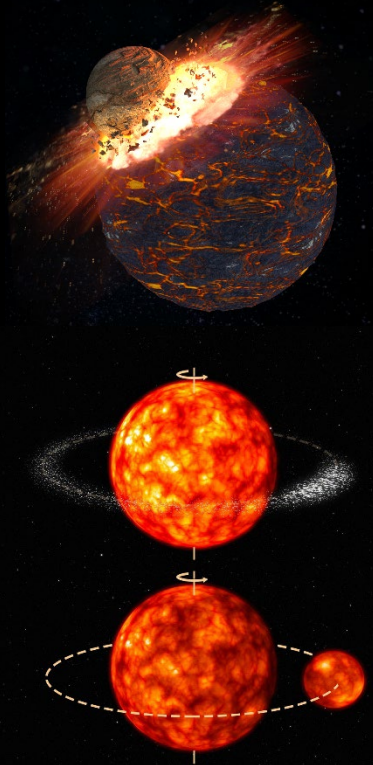


With an  $^{87}\text{Rb}/^{86}\text{Sr} = 0.005$  for Theia, you can barely reproduce present-day Earth  
(0.005 is the lowest measured volatile content for any meteorite!)

With an  $^{87}\text{Rb}/^{86}\text{Sr} = 0.01$  for Theia, you cannot reproduce present-day Earth

# Primary outcomes from Rb-Sr research

1. The precursor materials for Earth and Theia had to be volatile-poor (not a lot of water)
2. Nothing with a significant amount of Rb—such as all known asteroids and meteorites—could have been added after the Earth/Moon system formed



# There aren't that many options:

1. It was here from the start
2. It was delivered here
  - ~~By meteorites~~
  - By comets



# What about comets?

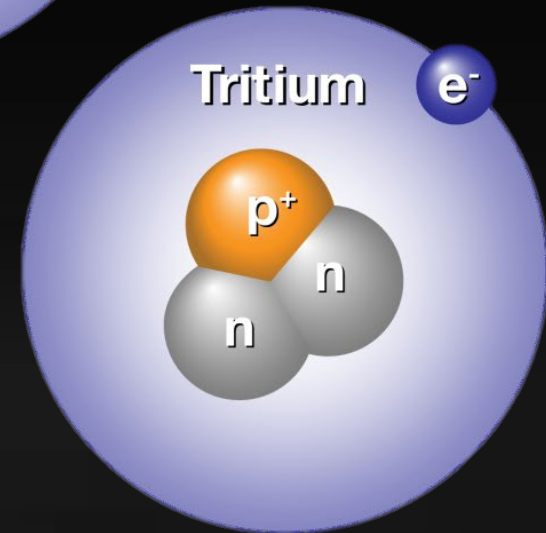
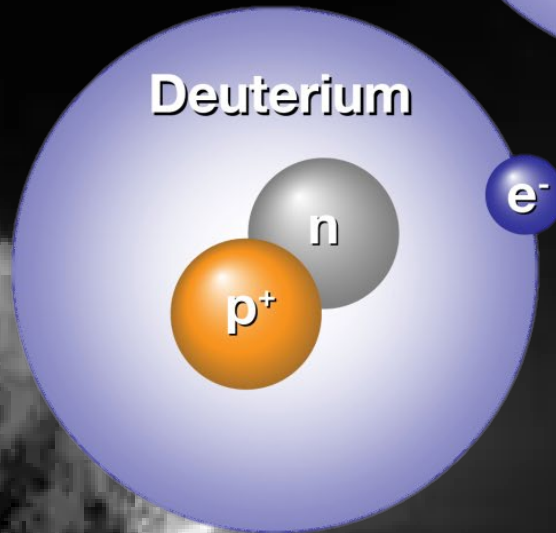
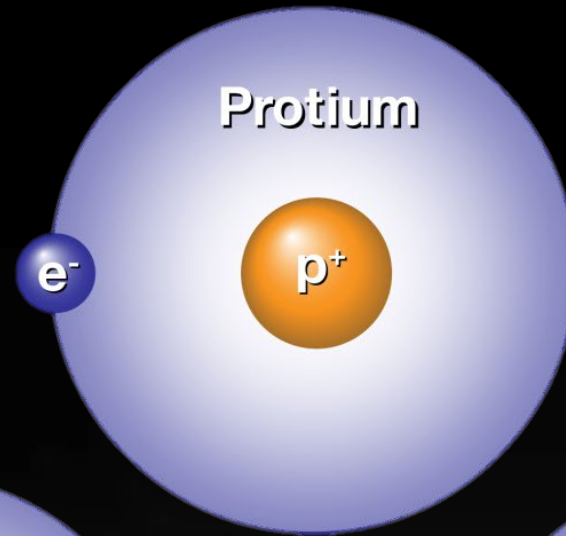
Isotopes might tell us!

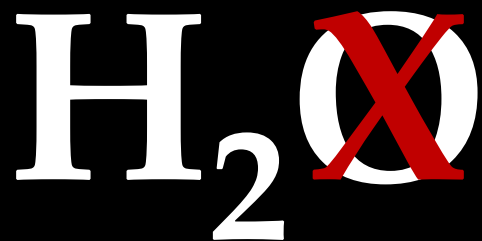


Hartley 2 – Credit NASA

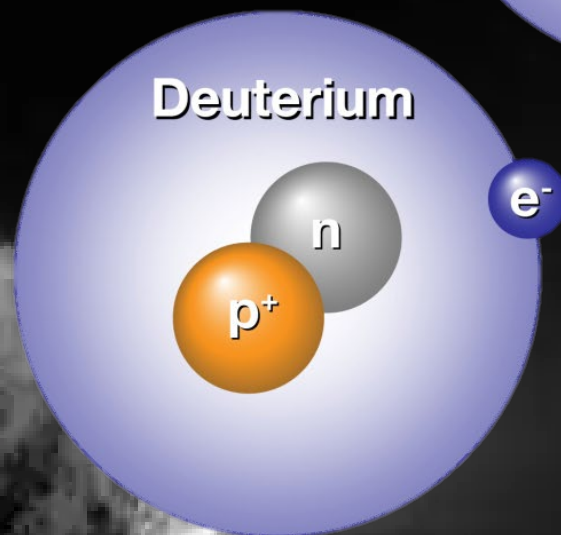
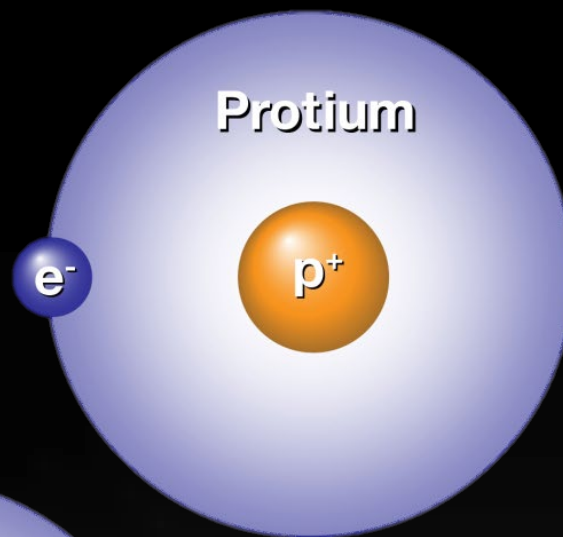
# $\text{H}_2\text{O}$

Isotopes might tell us!

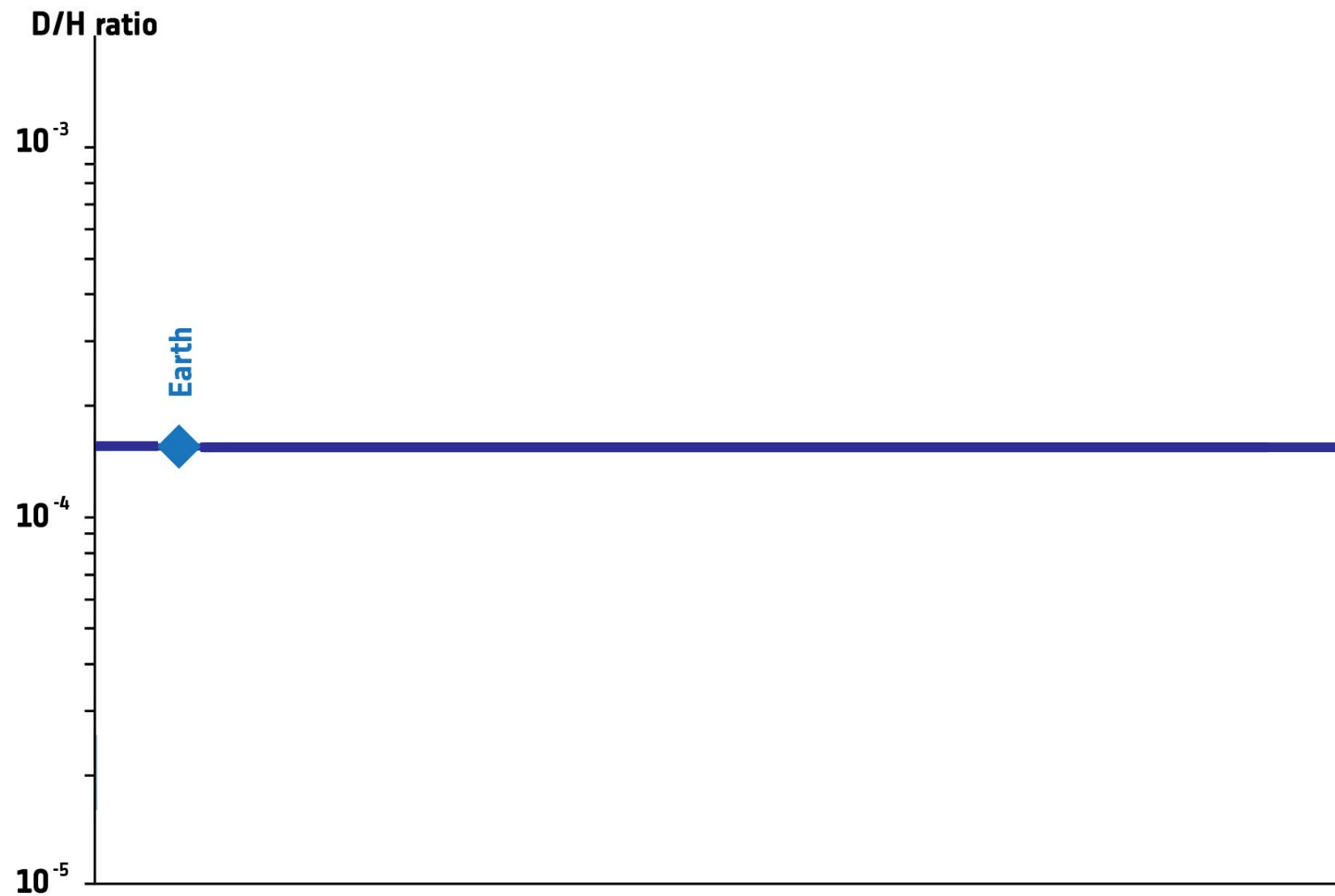
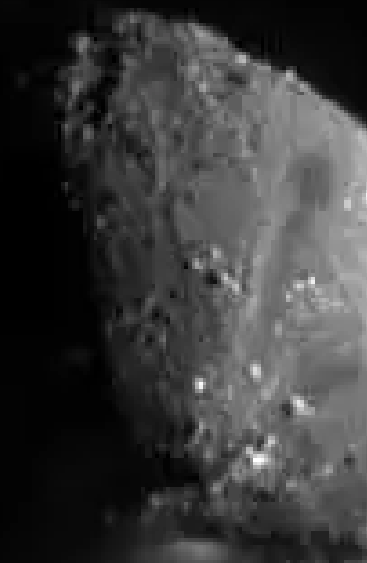


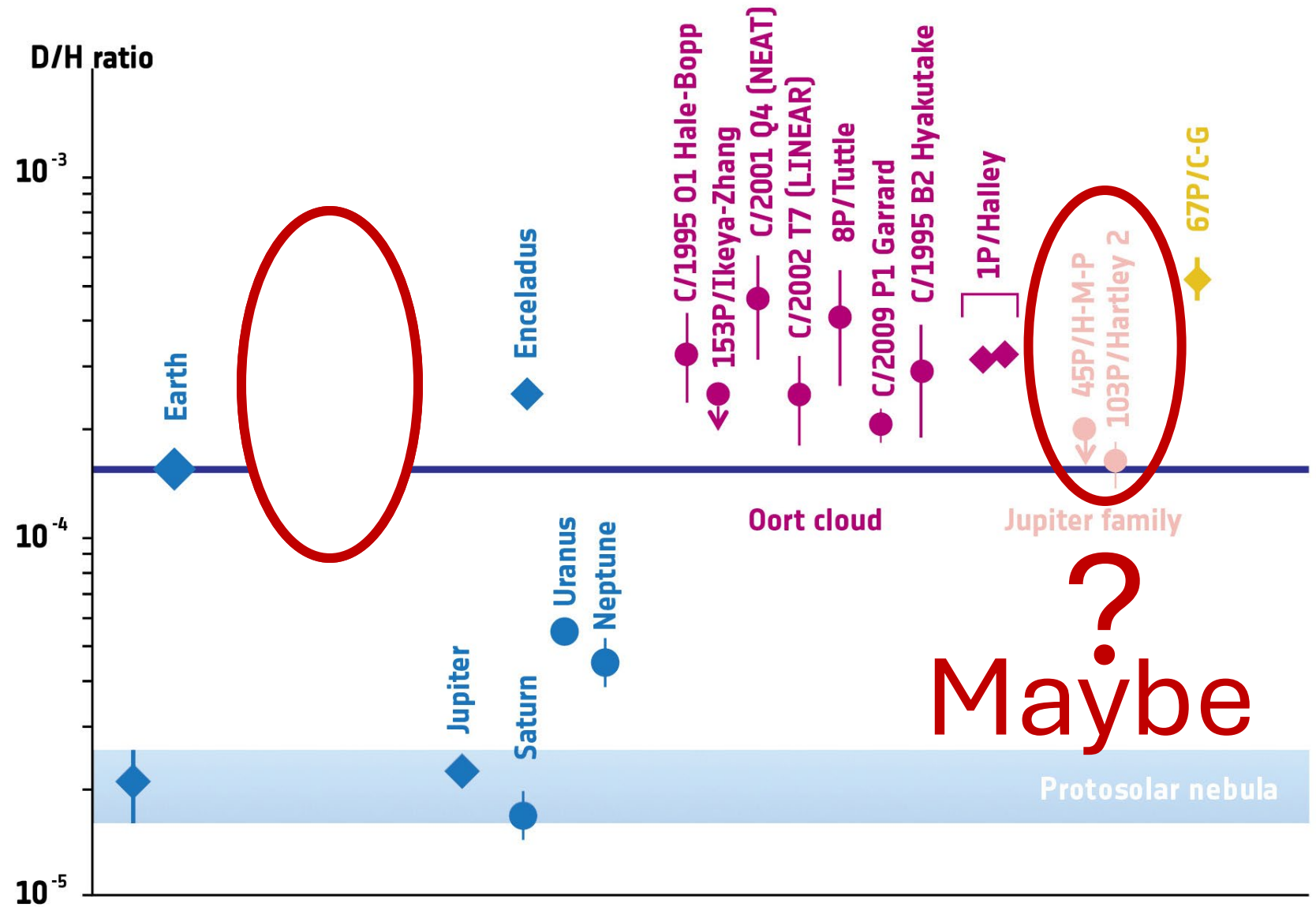


Isotopes might tell us!



# D/H





# What is the source of Earth's water?

1. It was here from the start?
2. It was delivered here
  - ~~By meteorites~~
  - By comets?



**Thoughts?  
Questions?**

